Surface Normal Integration for Convex Space-time Multi-view Reconstruction

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We propose a convex variational approach to space-time reconstruction which estimates surface normal information and integrates it into the photoconsistency estimation as well as into an anisotropic spatio-temporal total variation regularization. As such the proposed method generalizes the works [4], [5]. Although [4] already studied anisotropic regularization they did not estimate normals but used the normals from [2]. The combination of these methods, [4] and [2], is more than 40 times slower than our method as [4] alone needs about 1h to compute a single frame. In contrast, our method only takes about 3 minutes per frame including normal estimation and temporal regularization due to the proposed efficient implementation. Moreover, the method by Kolev et al. [4] does not work well on the 4D data sets we consider, as shown in [5, Fig. 5]. With the estimated normals at hand, we further propose an improvement of the photoconsistency voting scheme by Hernández and Schmitt [1] resulting in superior accuracy especially for sparse camera setups.

We represent the space-time surface as a binary interior/exterior labeling function $u: V \times T \mapsto \{0,1\}$ and state the spatio-temporal 3D reconstruction task as a minimization problem of the following energy.

$$E(u) = \int_{V \times T} \left[|\nabla_{\mathbf{x}} u|_{D_{\mathbf{x}}} + g_t |\nabla_t u| + \lambda f u \right] d\mathbf{x} dt$$
 (1)

The energy consists of three terms. An anisotropic spatial regularization term, defined be the norm $|\mathbf{y}|_{D_{\mathbf{x}}} = \langle \mathbf{y}, D_{\mathbf{x}}\mathbf{y} \rangle^{1/2}$ and the anisotropic diffusion matrix $D_{\mathbf{x}}(\mathbf{x},t) = \rho(\mathbf{x},t)^2 \mathbf{n} \mathbf{n}^T + \mathbf{n}_0 \mathbf{n}_0^T + \mathbf{n}_1 \mathbf{n}_1^T$ which lowers smoothing in the surface normal $\mathbf{n} \in \mathbb{R}^3$ direction and favors smoothness along the corresponding tangential directions \mathbf{n}_0 and $\mathbf{n}_1 = \mathbf{n} \times \mathbf{n}_0$. Further, the temporal regularization term weighted by function $g_t(\mathbf{x},t) = \exp\left(-a|\nabla f(\mathbf{x},t)|\right)$ accounts for a motion-dependent temporal smoothing. The purpose of the temporal regularization is to reduce surface jittering in scene parts with slow motion. Lastly, the data term, represented by function $f: V \times T \mapsto \mathbb{R}$ and smoothness weight λ , avoids trivial solutions of the energy and gives local preferences for an interior or exterior label. Both, the photoconsistency measure $\rho(\mathbf{x})$ in $D_{\mathbf{x}}$ and f depend on a voting scheme based on surface normal-dependent normalized cross-correlation (NCC) scores, represented by $C_i(\mathbf{x},d)$ for each point defined by the ray from camera i through point \mathbf{x} at distance d.

$$\rho(\mathbf{x}) = \exp\left[-\mu \sum_{i \in C'} \underbrace{\delta\left(d_i^{\max} = \operatorname{depth}_i(\mathbf{x})\right) \cdot C_i(\mathbf{x}, d_i^{\max})}_{\text{VOTE}_i(\mathbf{x})}\right]$$
(2)

The original voting scheme [1] computes the best depth hypothesis per camera ray as $d_i^{\max} = \arg\max_d C_i(\mathbf{x}, d)$ and does not enforce any spatial regularity of the votes, which we introduce by the following normal-dependent regularized voting scheme:

$$d_i^{\max} = \arg \max_{d} \int_{V_{\mathbf{x}}} C_i(\mathbf{x} - \mathbf{y}, d) \, \mathcal{G}(\mathbf{y}; \Sigma_{\mathbf{n}}) \, d\mathbf{y} , \qquad (3)$$

where $\mathcal{G}(y; \S{igma_n})$ is a normal-aligned anisotropic 3D Gaussian. We use surface normals at three places within our method: (a) NCC score, (b) voting scheme regularization and (c) anisotropic surface regularization. To estimate normals, we run our algorithm in two passes (see Fig. 1):

Pass 1: camera-to-point direction as normal for (a) and (b), isotropic surface regularization with high λ for (c)

Pass 2: normals from the previous pass for (a),(b) and (c) with lower λ for surface smoothness as desired

Finally, we propose an efficient GPU-accelerated primal-dual optimization of energy (1) which allows for comparatively low computation times. Our model yields significantly improved results over [5] which also compare well to other state-of-the-art reconstruction methods (see Fig. 2).

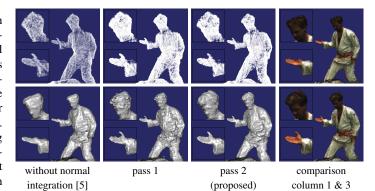


Figure 1: The photoconsistency measure $\rho(\mathbf{x})$ (top) and corresponding reconstructions (bottom) with and without the proposed normal integration. Pass 1 demonstrates the effect of the proposed normal-based voting regularization. Pass 2 adds anisotropic regularization and improved photoconsistency scores from pass 1.

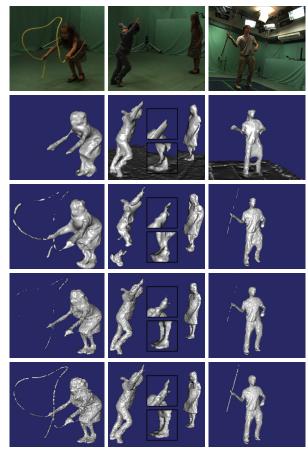


Figure 2: Reconstruction results of different methods. From top to bottom: input, Jancosek and Pajdla [3], PMVS+Poisson [2], Oswald and Cremers [5] and proposed. Our method recovers fine details and reduces temporal surface jittering.

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